

# Planets: Earth, Mars, & Beyond Research

# Aqueous Geochemistry and Astrobiology

- **Dr. Elisabeth (Libby) Hausrath**
- Professor
- Department of Geoscience
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## Expertise

- Using laboratory experiments, field work, and modeling to interpret water-rock interactions and soil-forming processes on Earth and Mars
- Interpreting the signatures of past aqueous and biological impacts on minerals
- Participating Scientist on the Mars Science Laboratory Curiosity and the Mars2020 rover Perseverance and member of the Network for Life Detection [\(NFOLD\)](#) Steering Committee..

# Holes made by sampling soil on Mars



Image credit: NASA/JPL-Caltech

<https://mars.nasa.gov/news/9311/nasas-perseverance-rover-gets-the-dirt-on-mars/#:~:text=The%20mission's%20first%20two%20samples,prepare%20for%20future%20missions%20there.>

# Geochemistry & Cosmochemistry

## Dr. Yan Hu

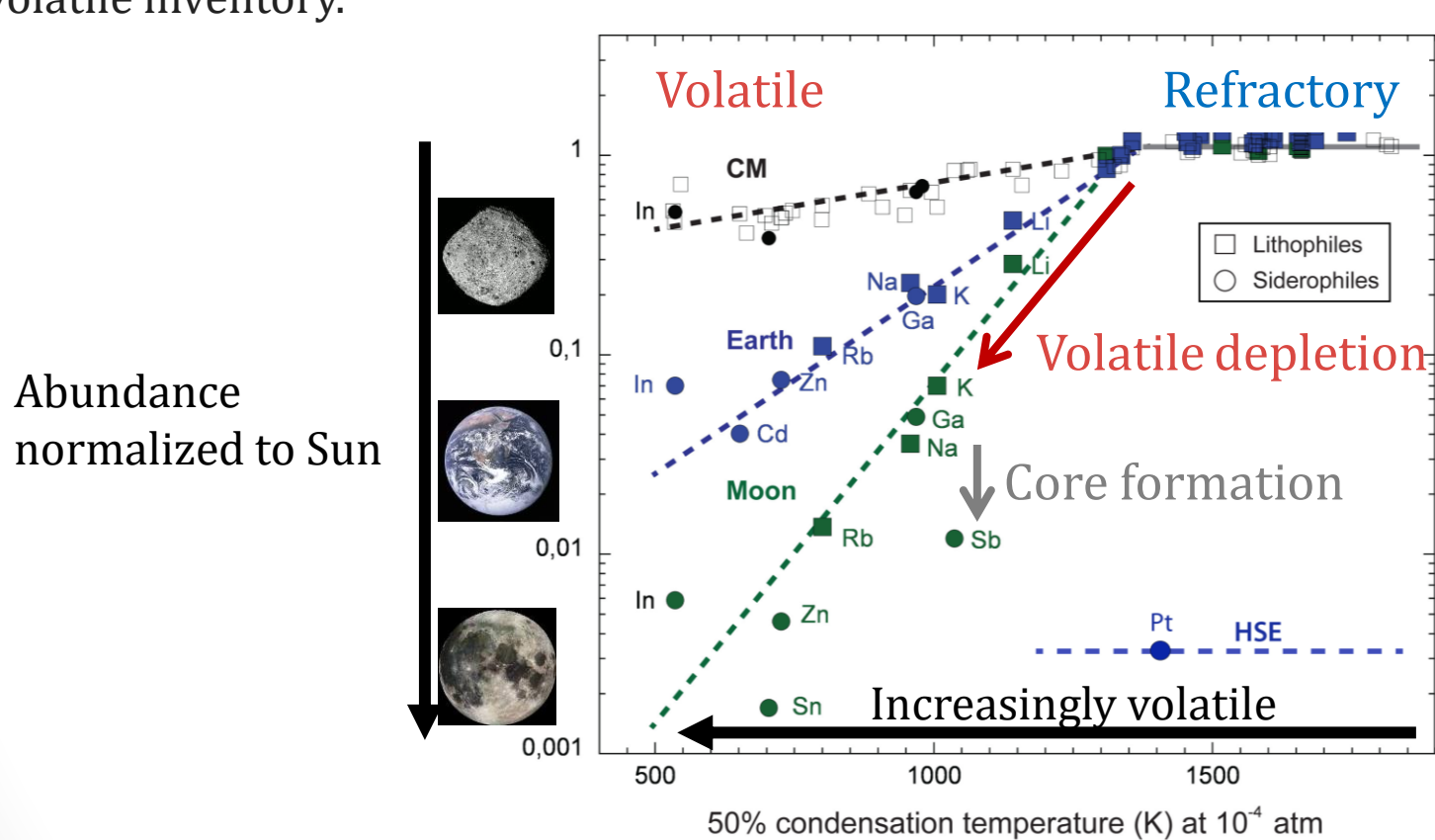
- Assistant Professor
- Department of Geoscience
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- Website: <https://sites.google.com/view/yan-hu>; [Google Scholar](#)

## Expertise

- Composition and evolution of Earth and early Solar System
- Subduction zone processes
- Non-traditional stable isotopes (e.g., Li, Mg, K, Ca, Fe)
- Multi-Collector Inductively Coupled Plasma Mass Spectrometry

# How planets formed and became habitable?

I study meteorites and mission-returned samples to learn about the building blocks of terrestrial planets and how condensation/evaporation shape their volatile inventory.



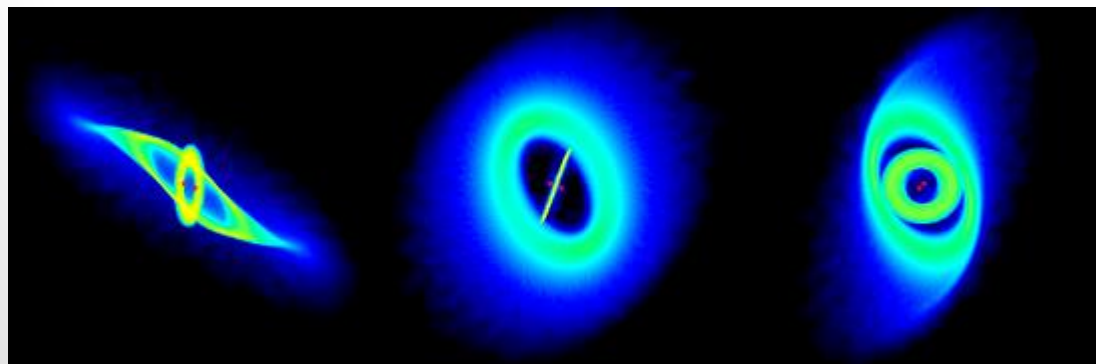
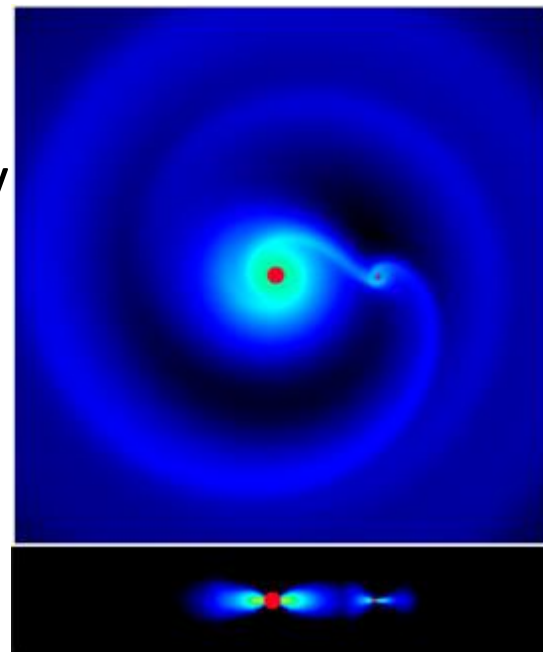
Volatile elements are variably depleted, making each planetary body unique.

# Rebecca Martin

- Professor of Astronomy
- BPB 233
- Department of Physics and Astronomy
- Email: [Rebecca.martin@unlv.edu](mailto:Rebecca.martin@unlv.edu)

## Expertise

- Star and planet formation
- Astrophysical Fluids
- Binary Star Systems
- Planetary System Dynamics



# Geomicrobiology

## **Dr. Aude Picard**

Assistant Research Professor

School of Life Sciences

[audeamelie.picard@unlv.edu](mailto:audeamelie.picard@unlv.edu)

## **Expertise**

- Anaerobic microbiology
- Microbial physiology
- Biomineralization
- Astrobiology and biosignatures
- Microscopy & spectroscopy

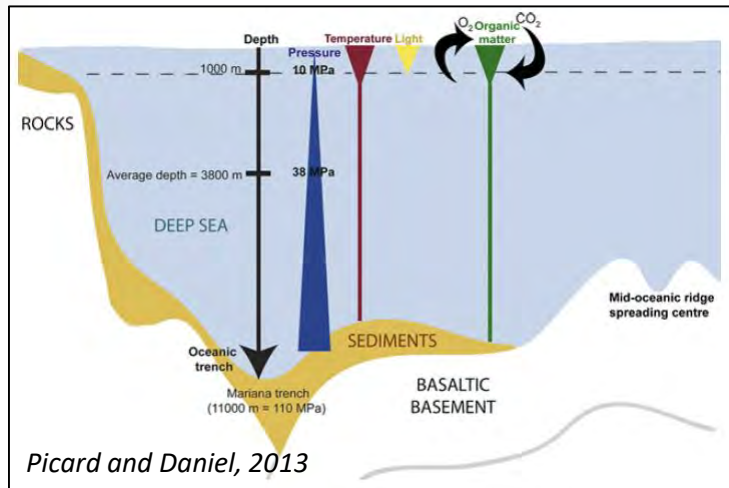


# Microbial life in extreme conditions

## ① Microbial life under high pressure

- What are the pressure limits for microbial life?

High-pressure environments represent the largest habitat for microbial life on Earth



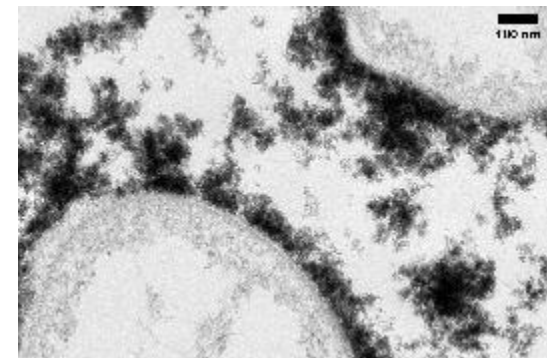
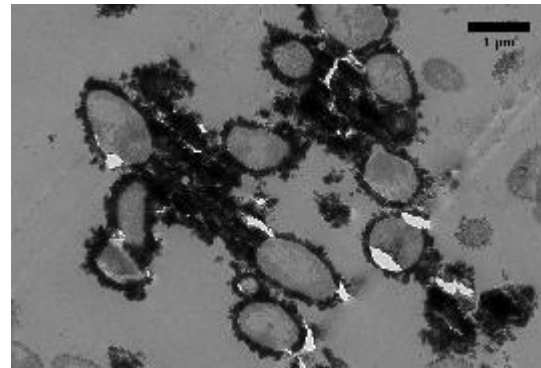
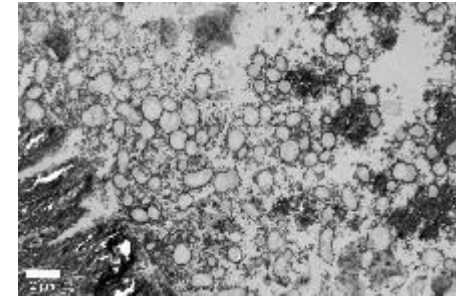
Oceans on icy moons (e.g. Europa) are potential habitats for microbial life in the outer Solar System



## ② Microbe-mineral interactions

- How do bacteria cope with mineral encrustation?
- Do minerals play a role in long-term survival of bacteria?

Transmission electron microscopy images of bacteria encrusted in iron sulfide minerals



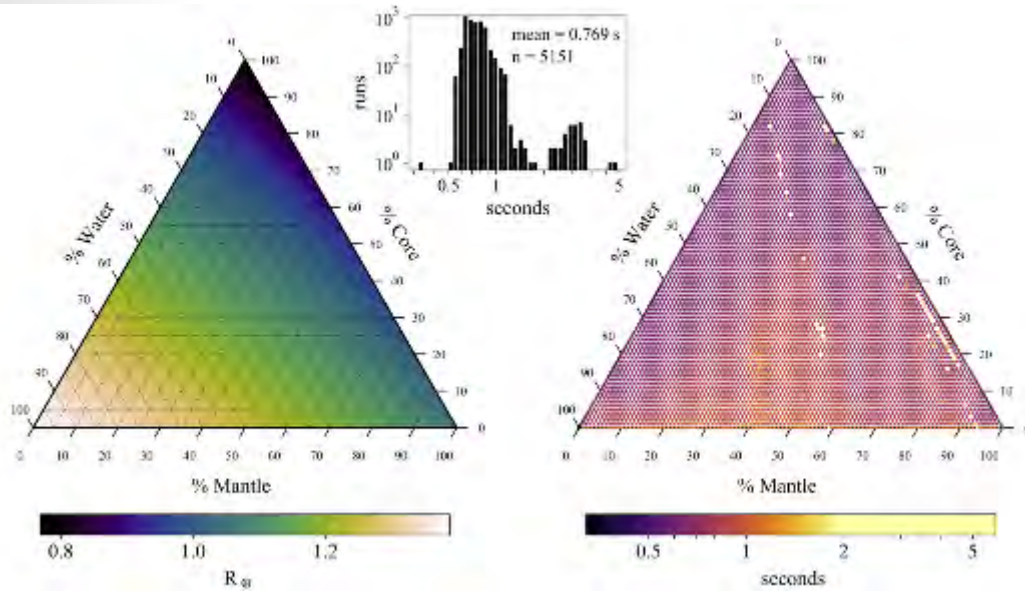


# Research Group of Dr. Steffen

- **Dr. Jason H. Steffen**
- Associate Professor
- Department of Physics and Astronomy
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- Website: [jasonhsteffen.com](http://jasonhsteffen.com)

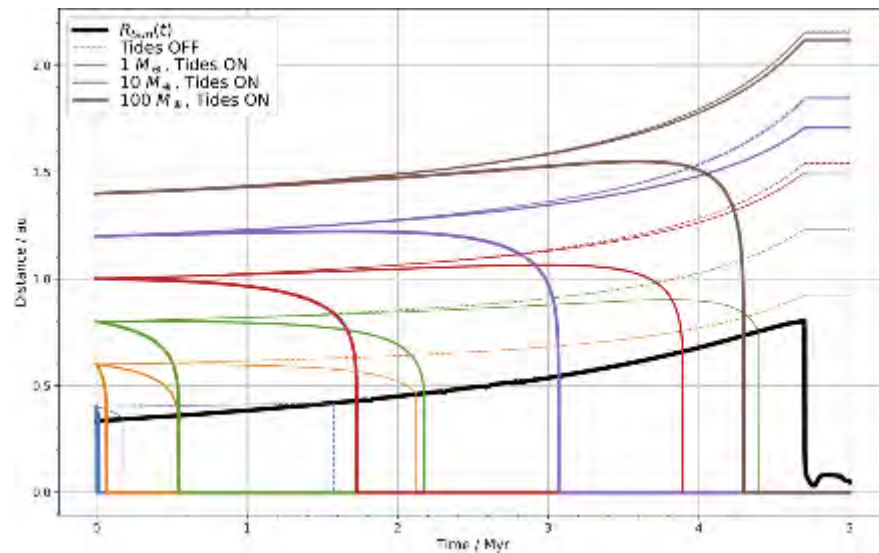
## Expertise

- Understanding the properties of extrasolar planets and planetary systems
- Planetary dynamics
- Planet interior modeling
- Composition of planet-forming materials



Timing results for planet models using the MAGRATHEA code, developed by our group at UNLV.

Future of planets in a system during the late stages of stellar evolution, including the effects of tides and stellar mass loss.



# Planetary petrology

## Dr. Arya Udry

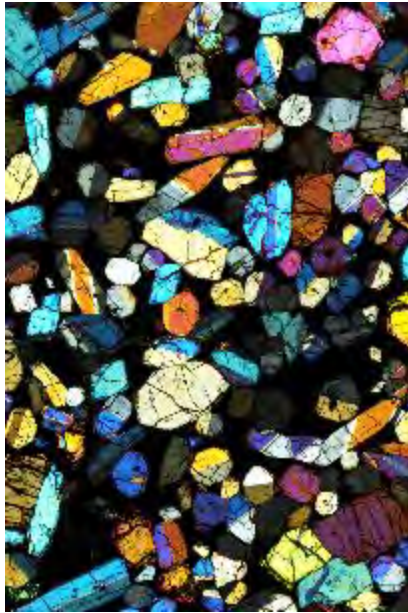
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## Expertise:

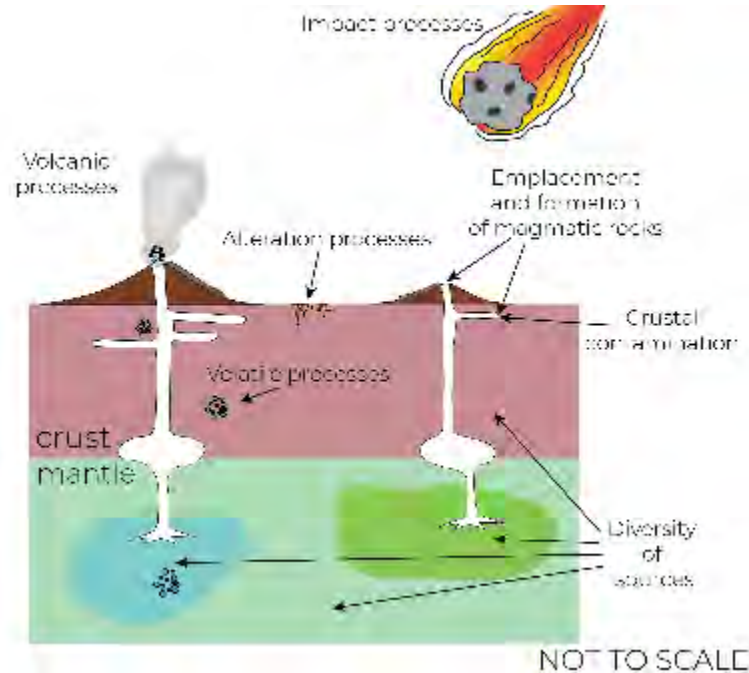
Planetary petrology

Martian igneous geology

# Martian geologic evolution using meteorites



*Polarized thin section image of nakhlite meteorite MIL 090030*



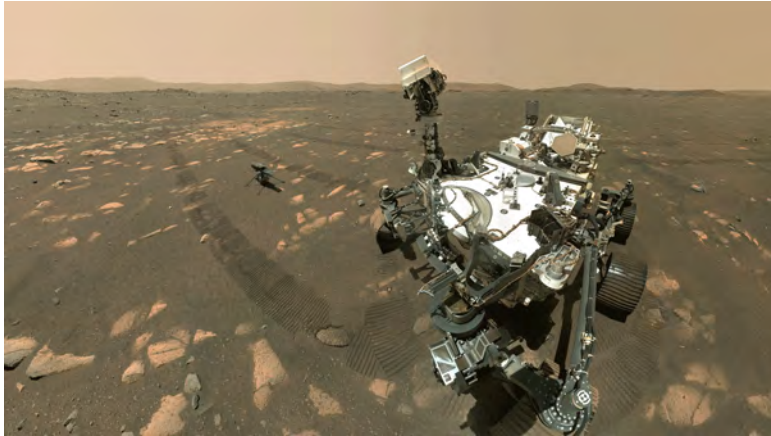
*Processes that can be understood using meteorites (Udry et al. 2020)*



*193 nm Excimer laser ablation system – Installed in 2021 to analyze mineral trace elements*

- I use meteorites, the only samples that we possess from Mars, to better constrain the interior composition and evolution of this planet
- Bulk rock and mineral geochemical down to the ppm scale

# Martian geologic evolution using rover analyses



Mars 2020 Perseverance and Ingenuity on Jezero crater – JPL/NASA image



**Early Mars (e.g., Noachian,  $\geq 3.7$  Ga?)**  
- Hotter, thinner crust  
- More crustal assimilation  
- **Enhanced magmatic evolution (more felsic and alkaline compositions)**  
\*not to scale

**Late Mars (e.g., Amazonian,  $\leq 3$  Ga?)**  
- Cooled, thickened, impacted crust (35-85 km average)<sup>1</sup>  
- Less crustal assimilation  
- **Less voluminous evolved magma**  
<sup>1</sup>Plesa et al., 2016

Models of magma on Mars (Ostwald et al., 2022)

- ❑ Thermodynamical modeling to understand formation of unique compositions of martian surface
- ❑ I am a participating scientist on the Mars2020 mission and I conduct modeling analyses to help understand the formation of magmatic rocks at Jezero crater

# Astrophysical Fluid Dynamics

## **Dr. Zhaohuan Zhu**

Department of Physics and Astronomy

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## **Expertise:**

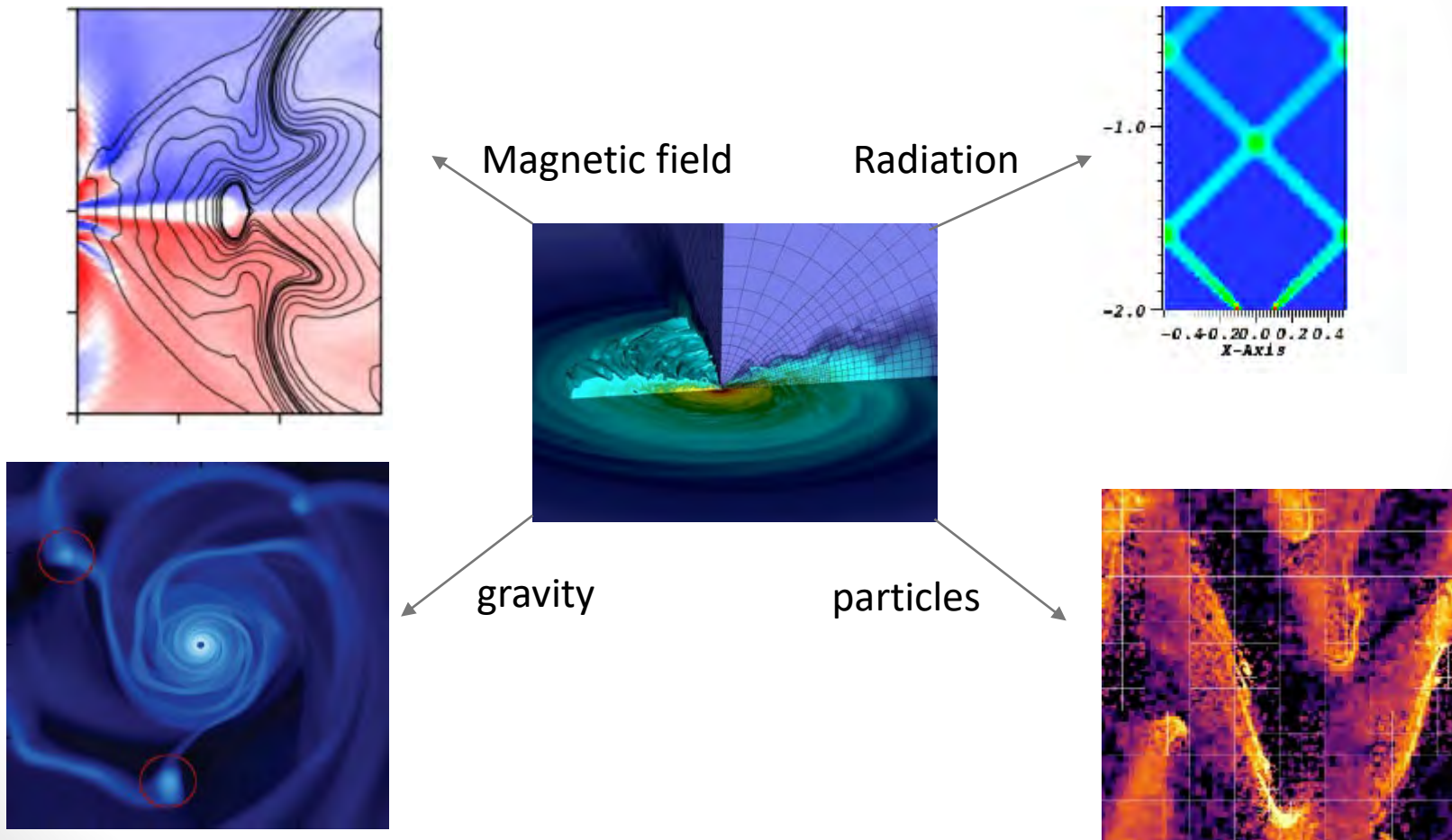
Fluid dynamics for astronomical project

Star and planet formation



# Fluid dynamics:

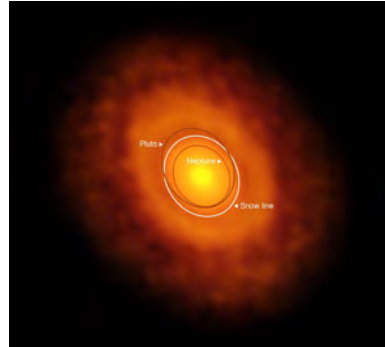
- Developing and using the state of the art numerical code to solve astrophysical fluid problem.





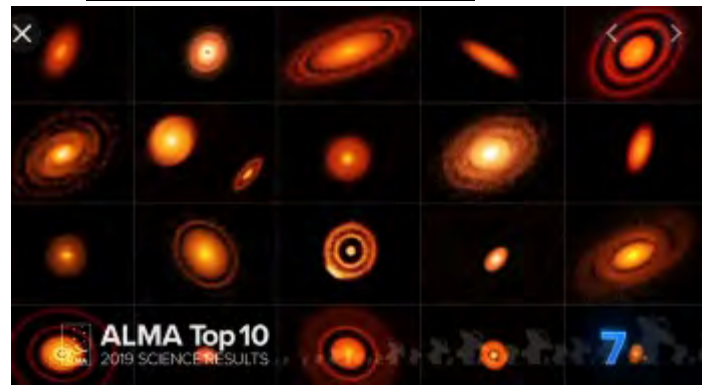
# Star and planet formation:

- Protoplanetary disk dynamics:



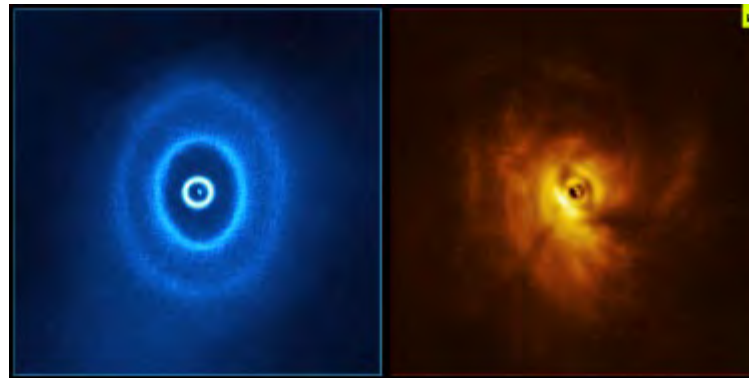
V883 Ori, *Nature*

- Planet formation



DSHARP

- Planet-disk interaction



GW Ori, *Science*